

# King's College London

UNIVERSITY OF LONDON

This paper is part of an examination of the College counting towards the award of a degree. Examinations are governed by the College Regulations under the authority of the Academic Board.

**B.Sc. EXAMINATION : Supplementary Assessment**

**CP/2620 Astrophysics**

**Summer 1999**

**It is the candidates' responsibility to ensure the safe delivery of their full written solutions to:**

**The Undergraduate Secretary, Dept. of Physics,  
King's College London, Strand, London WC2R 2LS**

**All work must be received by 13 September 1999 at the latest.**

**It is recommended that any postal submission uses Recorded Delivery.**

**Candidates should answer ALL parts of SECTION A,  
and ALL questions from SECTION B.**

**The approximate mark for each part of a question is indicated in square brackets.**

**TURN OVER WHEN INSTRUCTED  
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$$\begin{aligned}
\text{gravitational constant } G &= 6.673 \times 10^{-11} \text{ Nm}^2 \text{ kg}^{-2} . \\
\text{mass of the Sun } M_{\odot} &= 1.989 \times 10^{30} \text{ kg} . \\
\text{radius of the Sun } R_{\odot} &= 6.96 \times 10^8 \text{ m} . \\
\text{one parsec} &= 3.085 \times 10^{16} \text{ m} . \\
\text{speed of light } c &= 2.997 \times 10^8 \text{ m s}^{-1} . \\
\text{elementary charge } e &= 1.602 \times 10^{-19} \text{ C} . \\
\text{electron mass } m_e &= 9.109 \times 10^{-31} \text{ kg} . \\
\text{Planck constant } h &= 6.626 \times 10^{-34} \text{ J s} . \\
\text{Stefan-Boltzmann constant } \sigma &= 5.670 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4} .
\end{aligned}$$

## SECTION A – Answer ALL parts of this section

- 1.1) Define visual magnitude, absolute magnitude, colour index and luminosity. [7 marks]
- 1.2) Describe the method of trigonometrical parallax for finding the distances of stars. Why are results from the Hipparcos satellite important? The trigonometrical parallax of the Hyades cluster is measured as  $21.58 \times 10^{-3}$  arcseconds with an accuracy of  $0.12 \times 10^{-3}$  arcseconds. What is the distance and what is the error in the distance in parsecs? [7 marks]
- 1.3) Why are gamma-ray bursters believed to originate from neutron stars? [7 marks]
- 1.4) What observations have made reasonably certain that black holes exist of mass roughly  $10^9 M_{\odot}$  at the centre of some active Galaxies? [7 marks]
- 1.5) What is a black body? Why is a black body used as a standard source in radiation physics? A solitary neutron star that is not a pulsar is at a distance of 120 parsecs. The surface temperature is  $10^6$  K and the radius 20000 metres. What amount of energy per second from this star would fall on an area of 1 square meter just above the Earth's atmosphere? [7 marks]
- (                      Black body flux/m<sup>2</sup> =  $\sigma T^4$                       )
- 1.6) Why does the frequency of pulsars gradually decrease, and what physical parameter of the pulsar can be derived from the time constant of the decrease? [7 marks]

1.7) How may solar neutrinos be detected on the Earth's surface? Why is it important to detect them?

[7 marks]

1.8) Draw a diagram of a modern astronomical interferometer. What information may be obtained from such an interferometer?

[7 marks]

### SECTION B – Answer ALL questions

2) Describe the evolution of a star whose mass is  $5M_{\odot}$  from the start of the main sequence to the white dwarf stage.

[20 marks]

How may the mass of a star be estimated directly? What observations are required?

[5 marks]

Detailed analysis of two stars, masses  $M_1$  and  $M_2$ , that are in binary orbit shows that their distances from their center of mass are  $a_1 = 3.73 \times 10^{10}$  m and  $a_2 = 5.27 \times 10^{10}$  m. The period ( $P$ ) is 105 days. Find their masses.

[5 marks]

$$\left( \qquad (M_1 + M_2) = \frac{4\pi^2(a_1 + a_2)^3}{G P^2} \qquad \right)$$

3) What is a gamma-ray burster? Describe the observations that have been made. What evidence shows that some gamma-ray bursts come from distant galaxies.

[20 marks]

A gamma-ray burst from a source that is believed to be in a supernova remnant at a distance of 7 kiloparsec from the Earth produced a signal  $7 \times 10^{-6}$  J/m<sup>2</sup> in an Earth satellite. Estimate the total energy of the burst.

The source showed a periodicity of 5.16 seconds which is attributed to the rotation of a neutron star. Assuming a mass  $1.0M_{\odot}$  and a radius  $10^4$  metres for the neutron star, calculate the rotational kinetic energy. What fraction of the rotational energy is the burst energy?

[10 marks]

( Assume a homogeneous sphere with moment of inertia  $(2/5)MR^2$  )

- 4) Explain why Cepheid variable stars are used as distance indicators. What are the advantages and limitations of this method.

[20 marks]

Three Cepheid variables in the M100 galaxy in the Virgo cluster have periods  $P$  in days and mean *apparent* magnitudes  $m_v$  as follows.

$P$	$m_v$
53.1	24.8
39.5	25.4
30.1	26.4

The calibration of the mean *absolute* magnitude from the Hipparcos satellite is

$$\langle M_v \rangle = -1.43(\pm 0.1) - 2.81 \log_{10}(P(\text{days})).$$

Find the distance of each Cepheid variable in parsecs. If the M100 galaxy is receding from the Sun at  $1400 \pm 80$  kilometres/second, deduce the Hubble constant and estimate the error.

[10 marks]

$$(M_v - m_v = 5 - 5 \log_{10}(d))$$

- 5) Write an essay on pulsars.

[30 marks]