

King's College London

UNIVERSITY OF LONDON

This paper is part of an examination of the College counting towards the award of a degree. Examinations are governed by the College Regulations under the authority of the Academic Board.

B.Sc. EXAMINATION

CP/2620 Astrophysics

Summer 1998

Time allowed: THREE Hours

**Candidates should answer SIX parts of SECTION A,
and TWO questions from SECTION B.**

Separate answer books must be used for each Section of the paper.

The approximate mark for each part of a question is indicated in square brackets.

**You must not use your own calculator for this paper.
Where necessary, a College calculator will have been supplied.**

**TURN OVER WHEN INSTRUCTED
1998 ©King's College London**

$$\begin{aligned}
\text{gravitational constant } G &= 6.673 \times 10^{-11} \text{Nm}^2 \text{kg}^{-2} . \\
\text{mass of the Sun } M_{\odot} &= 1.989 \times 10^{30} \text{ kg} . \\
\text{radius of the Sun } R_{\odot} &= 6.96 \times 10^8 \text{ m} . \\
\text{one parsec} &= 3.085 \times 10^{16} \text{ m} . \\
\text{speed of light } c &= 2.997 \times 10^8 \text{ ms}^{-1} . \\
\text{elementary charge } e &= 1.602 \times 10^{-19} \text{ C} \\
\text{electron mass } m_e &= 9.109 \times 10^{-31} \text{ kg} . \\
\text{Planck constant } h &= 6.626 \times 10^{-34} \text{ Js} .
\end{aligned}$$

SECTION A – Answer SIX parts of this section

- 1.1) Define apparent magnitude, absolute magnitude, bolometric magnitude and luminosity. [7 marks]
- 1.2) Describe one method for finding the distance of a star in our Galaxy that does not depend on an auxiliary calibration of an astronomical scale. (The astronomical unit is assumed known.) [7 marks]
- 1.3) What are gamma-ray bursters? Discuss why are they of great interest in astrophysics? Show that an absorption feature at 20 keV which is due to cyclotron resonance, implies a magnetic field of about 2×10^8 Tesla. [7 marks]
- 1.4) What observations have made the existence reasonably certain of black holes of mass $\approx 2 - 8 M_{\odot}$ in our Galaxy? [7 marks]
- 1.5) Give reasons why astronomers think that Cepheid variable stars are stars with mass 5 to $10 M_{\odot}$ at a late stage in their lives. [7 marks]
- 1.6) Why are pulsars believed to be rotating neutron stars? [7 marks]
- 1.7) Suggest a reason why the solar neutrinos that are detected on the Earth's surface are less than the number which is predicted by the Standard Solar Model. [7 marks]

- 1.8) Discuss whether the time between the core collapse of a massive star and its explosion as a type II supernova can be estimated from observation. Can this time be predicted from theory?

[7 marks]

SECTION B – Answer TWO questions

- 2) What may be deduced from the observation of (a) visual binaries, (b) spectroscopic binaries and (c) eclipsing binaries?

[18 marks]

Two objects whose masses are M_1 and M_2 are in binary orbit. Assuming circular orbits and Newtonian gravity, show that

$$\omega^2 = \frac{G(M_1 + M_2)}{(a_1 + a_2)^3}$$

where a_1 and a_2 are the radii of the orbits about the centre of mass and ω is the angular velocity.

The first object is observed spectroscopically and has an apparent velocity V with respect to the centre of mass. The angle between the plane of the orbit and the plane of the sky is i . From this and $M_1 a_1 = M_2 a_2$ deduce that

$$\frac{(M_2/M_1)^3 (\sin i)^3}{(1 + M_2/M_1)^2} = \frac{\omega^2 (V/\omega)^3}{G M_1}$$

An x-ray source has an orbiting star whose period is 4.10 days and sinusoidal orbital velocity about 120 km/sec with respect to the centre of mass. Spectroscopic data suggest the star is a B0 supergiant whose mass, by comparison with other such stars, is estimated to be $\approx 9M_\odot$. The light is not eclipsed. The light from the other object is a blue glow with no absorption lines visible. What can be deduced about the mass and nature of this other object?

[12 marks]

- 3) Describe the evolution of a star of mass $\approx 5M_\odot$ according to current theories from its first appearance on the main sequence to its final state.

[30 marks]

- 4) Describe the observed properties of Cepheid variable stars. Explain how they are used as distance indicators. State the advantages and limitations of this method.

[30 marks]

- 5) Why do some stars at the end of their lives, explode as supernovas? What is known about supernovas from observation?

[30 marks]