

RELATIVITY (MTH6132)

SOLUTIONS TO THE PROBLEM SET 9

1. To compute the timelike geodesic equations recall that

$$L = - \left(1 - \frac{2GM}{r}\right) \dot{t}^2 + \left(1 - \frac{2GM}{r}\right)^{-1} \dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \dot{\varphi}^2$$

The geodesic equations are then obtained by noticing that

$$\frac{\partial L}{\partial t} = 0, \quad \frac{\partial L}{\partial \varphi} = 0, \quad \frac{\partial L}{\partial r} = 2r(\dot{\theta}^2 + \sin^2 \theta \dot{\varphi}^2) - A^{-2} A' \dot{r}^2 - A' \dot{t}^2, \quad \frac{\partial L}{\partial \theta} = r^2 \sin \theta \cos \theta \dot{\varphi}^2,$$

where

$$A = 1 - \frac{2GM}{r}, \quad A' = \frac{2GM}{r^2},$$

and ' denotes the derivative with respect to r . The actual expressions are then

$$\frac{d}{d\tau} (A\dot{t}) = 0, \tag{1}$$

$$\frac{d}{d\tau} (2A^{-1}\dot{r}) - 2r(\dot{\theta}^2 + \sin^2 \theta \dot{\varphi}^2) + A^{-2} A' \dot{r}^2 + A' \dot{t}^2 = 0, \tag{2}$$

$$\frac{d}{d\tau} (r^2 \dot{\theta}) - r^2 \sin \theta \cos \theta \dot{\varphi}^2 = 0, \tag{3}$$

$$\frac{d}{d\tau} (r^2 \sin^2 \theta \dot{\varphi}) = 0. \tag{4}$$

In addition one has equation (5).

2. For a timelike geodesic one has that

$$\left(1 - \frac{2GM}{r}\right) \dot{t}^2 - \left(1 - \frac{2GM}{r}\right)^{-1} \dot{r}^2 - r^2 \dot{\theta}^2 - r^2 \sin^2 \theta \dot{\varphi}^2 = 1. \tag{5}$$

A radial geodesic is defined by the conditions

$$\dot{\theta} = \dot{\varphi} = 0.$$

Now, the geodesic equation for the time coordinate gives

$$\dot{t} = l \left(1 - \frac{2GM}{r}\right)^{-1}$$

Substituting this into equation (5) one obtains

$$l^2 \left(1 - \frac{2GM}{r}\right)^{-1} - 1 = \dot{r}^2 \left(1 - \frac{2GM}{r}\right)^{-1},$$

so that

$$\dot{r}^2 = l^2 - 1 + \frac{2GM}{r}.$$

If $l = 1$, the latter gives

$$\dot{r} = \frac{dr}{d\tau} = \sqrt{\frac{2GM}{r}}.$$

Thus,

$$\int_0^{2GM} r^{1/2} dr = \sqrt{2GM} \int_{\tau_1}^{\tau_2} = \sqrt{2GM}(\tau_2 - \tau_1) = \sqrt{2GM}\Delta\tau,$$

from where

$$\frac{2}{3}(2GM)^{3/2} = (2GM)^{1/2}\Delta\tau.$$

Hence

$$\Delta\tau = \frac{4GM}{3}.$$

3. Now consider an orbit in the Equatorial plane ($\theta = \pi/2$). Hence $\dot{\theta} = 0$. Furthermore, the orbit is circular so that $r = D$, and $\dot{r} = 0$. The integration of equation (1) gives

$$\dot{t} = l \left(1 - \frac{2GM}{r}\right)^{-1} = l \left(1 - \frac{2GM}{D}\right)^{-1}, \quad l \text{ a constant.}$$

Also, from equation 4 one has

$$\frac{d}{d\tau} (r^2\dot{\varphi}) = 0 \Rightarrow r^2\dot{\varphi} = h, \quad \text{a constant.}$$

Notice that equation (2) reduces to

$$-2r\dot{\varphi}^2 + A'\dot{t}^2 = 0$$

Thus, substituting the previous two equations one has that

$$-2\frac{h^2}{D^3} + l^2 \left(1 - \frac{2GM}{D}\right)^{-2} \frac{2GM}{D^2} = 0 \Rightarrow l^2 = \left(1 - \frac{2GM}{D}\right)^2 \frac{h^2}{GMD}. \quad (6)$$

Now, for this problem, equation (5) gives

$$1 = \left(1 - \frac{2GM}{D}\right) \dot{t}^2 - D^2\dot{\varphi}^2 = 1 = l^2 \left(1 - \frac{2GM}{D}\right)^{-1} - \frac{h^2}{D^2}.$$

Substituting (6) into the last equation one obtains

$$\left(1 - \frac{2GM}{D}\right) \frac{h^2}{GMD} - \frac{h^2}{D^2} = 1 \Rightarrow h^2 = GMD \left(1 - \frac{3GM}{D}\right)^{-1}.$$

Finally, following the hint, the time elapsed in one orbit is given by

$$\tau = \int_0^{2\pi} \frac{d\tau}{d\varphi} d\varphi = \int_0^{2\pi} \dot{\varphi}^{-1} d\varphi = \int_0^{2\pi} \frac{D^2}{h} d\varphi = \frac{2\pi D^2}{h},$$

where equation (4) has been used. Substituting the obtained value for h one gets

$$\tau = 2\pi \left(\frac{D^3}{GM}\right)^{1/2} \left(1 - \frac{3GM}{D}\right)^{1/2},$$

as required.