

RELATIVITY (MTH6132)

SOLUTIONS TO THE PROBLEM SET 8

1. In this case one has that

$$L = -e^{2Ar} \dot{t}^2 + \dot{r}^2.$$

Now, for the t components one has that

$$\frac{\partial L}{\partial t} = 0, \quad \frac{\partial L}{\partial \dot{t}} = -2e^{2Ar} \dot{t}, \quad \frac{d}{d\lambda} (-2e^{2Ar} \dot{t}) = \ddot{t} e^{2Ar} + 2Ae^{2Ar} \dot{r} \dot{t} = 0.$$

Thus, the Euler-Lagrange equation is given by

$$\ddot{t} + 2A\dot{r}\dot{t} = 0.$$

Now, comparing with the geodesic equation

$$\ddot{x}^a + \Gamma^a_{bc} \dot{x}^b \dot{x}^c = \ddot{t} + \Gamma^t_{tt} \dot{t}^2 + 2\Gamma^t_{tr} \dot{t} \dot{r} + \Gamma^t_{rr} \dot{r}^2 = 0.$$

Comparing one gets

$$\Gamma^t_{tt} = 0, \quad \Gamma^t_{tr} = \Gamma^t_{rt} = A, \quad \Gamma^t_{rr} = 0.$$

For the r components one has that

$$\frac{\partial L}{\partial r} = -2Ae^{2Ar} \dot{t}^2, \quad \frac{\partial L}{\partial \dot{r}} = 2\dot{r}, \quad \frac{d}{d\lambda} \left(\frac{\partial L}{\partial \dot{r}} \right) = 2\ddot{r}.$$

Hence, the Euler-Lagrange equation is given by

$$\ddot{r} + Ae^{2Ar} \dot{t}^2 = 0.$$

Again, comparison with the geodesic equation gives

$$\Gamma^r_{rr} = 0, \quad \Gamma^r_{rt} = \Gamma^r_{tr} = 0, \quad \Gamma^r_{tt} = Ae^{2Ar}.$$

Using the formula for the Ricci tensor one has that

$$\begin{aligned} R_{tt} = R_{00} &= \partial_a \Gamma^a_{00} - \partial_0 \Gamma^a_{0a} + \Gamma^a_{ea} \Gamma^e_{00} - \Gamma^a_{e0} \Gamma^e_{0a}, \\ &= \partial_0 \Gamma^0_{00} + \partial_1 \Gamma^1_{00} - \partial_0 \Gamma^0_{00} - \partial_0 \Gamma^1_{01} + \Gamma^0_{e0} \Gamma^e_{00} + \Gamma^1_{e1} \Gamma^e_{00} - \Gamma^0_{e0} \Gamma^e_{00} - \Gamma^1_{e0} \Gamma^e_{01}, \\ &= \partial_1 \Gamma^1_{00} + \Gamma^0_{e0} \Gamma^e_{00} + \Gamma^1_{e1} \Gamma^e_{00} - \Gamma^0_{e0} \Gamma^e_{00} - \Gamma^1_{e0} \Gamma^e_{01}, \\ &= \partial_1 \Gamma^1_{00} + \Gamma^0_{00} \Gamma^0_{00} + \Gamma^0_{10} \Gamma^1_{00} + \Gamma^1_{01} \Gamma^0_{00} + \Gamma^1_{11} \Gamma^1_{00} - \Gamma^0_{00} \Gamma^0_{00} - \Gamma^0_{10} \Gamma^1_{00} - \Gamma^1_{00} \Gamma^0_{01} \\ &\quad - \Gamma^1_{10} \Gamma^1_{01}, \\ &= \partial_1 \Gamma^1_{00} + \Gamma^0_{10} \Gamma^1_{00} - \Gamma^0_{10} \Gamma^1_{00} - \Gamma^1_{00} \Gamma^0_{01}, \\ &= A \partial_r (e^{2Ar}) - A^2 e^{2Ar} = A^2 e^{2Ar}. \end{aligned}$$

2. (i) Starting from the definition of the Christoffel symbols:

$$\begin{aligned} \Gamma_{abc} &\equiv g_{af} \Gamma^f_{bc} \\ &= g_{af} \Gamma^f_{bc} = \frac{1}{2} g_{af} g^{fe} (\partial_b g_{ec} + \partial_c g_{be} - \partial_e g_{bc}) \\ &= \frac{1}{2} \delta_a^e (\partial_b g_{ec} + \partial_c g_{be} - \partial_e g_{bc}) \\ &= \frac{1}{2} (\partial_b (\delta_a^e g_{ec}) + \partial_c (\delta_a^e g_{be}) - \partial_a g_{bc}) \\ &= \frac{1}{2} (\partial_b g_{ac} + \partial_c g_{ba} - \partial_a g_{bc}). \end{aligned}$$

(ii) One can verify this by direct inspection. For example,

$$R_{bacd} = K(g_{bc}g_{ad} - g_{bd}g_{ca}) = -K(g_{ac}g_{bd} - g_{ad}g_{cb}) = -R_{abcd},$$

where it has been used that $g_{ab} = g_{ba}$. To compute $\nabla_e R_{abcd}$ one uses the Leibnitz rule:

$$\nabla_e R_{abcd} = K((\nabla_e g_{ac})g_{bd} + g_{ac}(\nabla_e g_{bd}) - (\nabla_e g_{ac})g_{bd} - g_{ac}(\nabla_e g_{bd})) = 0.$$

To compute R_{bd} proceed as follows:

$$\begin{aligned} R_{bd} &= g^{ac} R_{abcd} = K g^{ac} (g_{ac}g_{bd} - g_{ad}g_{cb}) = K (g^{ac} g_{ac}g_{bd} - g^{ac} g_{ad}g_{cb}) \\ &= K (\delta_a^a g_{bd} - \delta_d^c g_{cb}) \\ &= K (4g_{bd} - g_{bd}) = 3K g_{bd}, \end{aligned}$$

where it has been used that $\delta_a^a = 4$ —see Coursework 6. Finally

$$R = g^{bd} R_{bd} = 3K g^{bd} g_{bd} = 12K.$$

3. (i) Start from the expression for the Riemann tensor in locally inertial coordinates:

$$R_{abcd} = \frac{1}{2} (\partial_d \partial_a g_{bc} + \partial_c \partial_b g_{ad} - \partial_c \partial_a g_{bd} - \partial_d \partial_b g_{ac}).$$

Now, performing the substitutions $b \rightarrow c \rightarrow d \rightarrow b$ twice one obtains

$$\begin{aligned} R_{acdb} &= \frac{1}{2} (\partial_b \partial_a g_{cb} + \partial_d \partial_c g_{ab} - \partial_d \partial_a g_{cb} - \partial_b \partial_c g_{ad}), \\ R_{adb c} &= \frac{1}{2} (\partial_c \partial_a g_{dc} + \partial_b \partial_d g_{ac} - \partial_b \partial_a g_{dc} - \partial_c \partial_d g_{ab}). \end{aligned}$$

Now, adding and recalling that partial derivatives commute and that g_{ab} is symmetric one obtains the desired result.

(ii) There is a typo in this expression and it should read

$$R_{a[bcd]} = 0.$$

Note that (see notes on symmetric and antisymmetric parts of a tensor):

$$R_{a[bcd]} = \frac{1}{6} (R_{abcd} + R_{acdb} + R_{adb c} - R_{acbd} - R_{adcb} - R_{abd c}).$$

Recall, however, that the Riemann tensor is antisymmetric under the interchange of the last two indices —e.g. $R_{abcd} = -R_{abdc}$. Thus,

$$R_{a[bcd]} = \frac{1}{3} (R_{abcd} + R_{acdb} + R_{adb c}) = 0.$$

Also notice that the wrong expression

$$R_{a(bcd)} = 0$$

is also true. This again follows from the antisymmetry of the Riemann tensor on the last two indices:

$$R_{abcd} = -R_{abdc} \Rightarrow R_{a(bcd)} = -R_{a(bdc)} = -R_{a(bcd)},$$

from where it follows that

$$R_{a(bcd)} = 0.$$