

# SOLAR SYSTEM

①

## EXERCISE SHEET 3

### SOLUTIONS

$$\boxed{Q3.1} \quad h = na^2 \sqrt{1-e^2}$$

But  $h = r^2 \dot{\theta}$  where  $\dot{\theta}$  is angular velocity  
(H.B.  $n = \dot{\theta}$  in circular case)

$$\text{also } r = \frac{a(1-e^2)}{1+e \cos f} \quad (5)$$

$$\text{i.e. } r^2 = \frac{a^2(1-e^2)^2}{(1+e \cos f)^2}; \quad \dot{\theta} = \frac{h}{r^2} = \frac{na^2 \sqrt{1-e^2} (1+e \cos f)^2}{a^2(1-e^2)^2}$$

$$\text{i.e. } \dot{\theta} = \frac{n \sqrt{1-e^2} (1+e \cos f)^2}{(1-e^2)^2} \quad (10)$$

$$\text{at aphelion, } \dot{\theta} = \frac{n \sqrt{1-e^2} (1-e)^2}{(1-e^2)^2}$$

(i.e.  $f = \pi$ )

$$= \frac{n \sqrt{1-e^2} (1-e)^2}{(1+e)^2 (1-e)^2} = \frac{n \sqrt{1-e^2}}{(1+e)^2} \quad (10)$$

Jupiter's mean motion,  $n' = \dot{\theta}'$  (a constant, as it is a circular orbit).

$$\text{at the 2:1 resonance, } n = 2n' \quad (5)$$

Matching the angular velocity at aphelion gives

$$\dot{\theta} = \frac{n\sqrt{1-e^2}}{(1+e)^2} = n' = \frac{n}{2}$$

i.e. asteroid would be instantaneously stationary when  $e$  satisfies the equation

$$\frac{\sqrt{1-e^2}}{(1+e)^2} = \frac{1}{2}$$

$$\text{i.e. } 4(1-e^2) = (1+e)^4$$

$$\text{i.e. } 4(1+e)(1-e) = (1+e)^4$$

$$\text{i.e. } 4(1-e) = (1+e)^3$$

$$\text{or } 4 - 4e = 1 + 3e + 3e^2 + e^3$$

$$\text{i.e. } e^3 + 3e^2 + 7e - 3 = 0$$

The solutions are :  $-1.68233 \pm 2.32308i$

$$\text{or } 0.364656$$

$e$  must be real (and +ve), so answer is

$$e = 0.364656$$

(2)

5

10

5

Q3.2

$$\dot{\bar{\omega}} = \frac{1}{na^2 e} \frac{\partial R}{\partial \bar{\omega}}$$

(3)

$$R = \frac{Gm'}{a'} \left\{ \frac{A}{8} (e^2 + e'^2) + \frac{B}{4} ee' \cos(\bar{\omega}' - \bar{\omega}) \right\}$$

$$\text{i.e. } \dot{\bar{\omega}} = \frac{Gm'}{a'} \cdot \frac{1}{na^2 e} \left\{ \frac{A}{8} \cdot 2e + \frac{B}{4} e' \cos(\bar{\omega}' - \bar{\omega}) \right\} \quad (10)$$

Use  $GM \approx n^2 a^3$  which gives

$$n \approx \frac{360^\circ}{a^{3/2}} = \frac{360}{(2.86)^{3/2}} = 74.43 \text{ }^\circ/\text{yr} = 7443 \text{ }^\circ/\text{cent.} \quad (10)$$

$$\text{and } \dot{\bar{\omega}} = \frac{n^2 a^3}{a'} \frac{m'}{M} \cdot \frac{1}{na^2} \left\{ \frac{A}{4} + \frac{B}{4} \frac{e'}{e} \cos(\bar{\omega}' - \bar{\omega}) \right\}$$

$$= \frac{n}{4} \frac{a}{a'} \frac{m'}{M} \left\{ A + B \frac{e'}{e} \cos(\bar{\omega}' - \bar{\omega}) \right\}$$

$$= \frac{7443}{4} \cdot \frac{2.86}{5.20} \cdot 10^{-3} \left\{ 1.79185 + \frac{0.048}{0.15} (-1.18029) \times \cos(\bar{\omega}' - \bar{\omega}) \right\}$$

$$= 1.0234 \left\{ 1.79185 - 0.37769 \cos(\bar{\omega}' - \bar{\omega}) \right\} \quad (10)$$

$$\text{i.e. } \left( \dot{\bar{\omega}} \right)_{\bar{\omega}' = \bar{\omega}} = 1.44725 \text{ }^\circ/\text{cent} \quad (5)$$

$$\left( \dot{\bar{\omega}} \right)_{\bar{\omega}' = \bar{\omega} + 180^\circ} = 2.22031 \text{ }^\circ/\text{cent.} \quad (5)$$

Therefore the precession rate is slower when the perihelia are aligned. Therefore, at any given time, there should be more asteroids with perihelia aligned with Jupiter.

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