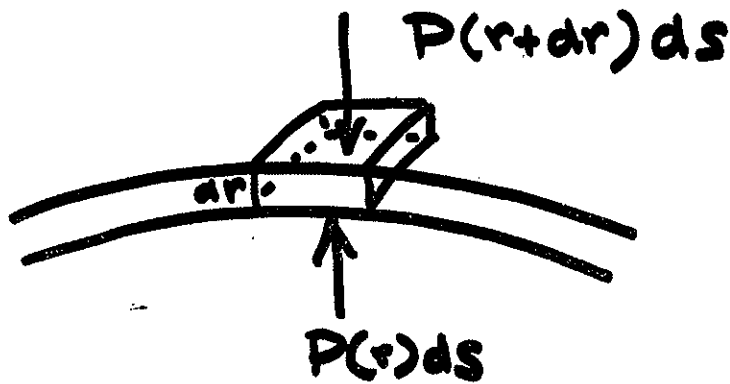


## Hydrostatic equilibrium

(3)

Consider an infinitesimal volume element between  $r$  and  $r+dr$  with base  $ds$



If  $P(r)$  is the total pressure on the bottom area and  $P(r+dr)$  is the pressure on the top. The buoyant pressure force in the direction of increasing  $r$  is

$$\begin{aligned} & P(r)ds - P(r+dr)ds \\ &= P(r)ds - \left( P(r)ds + \frac{dP}{dr} dr ds \right) \\ &= -\frac{dP}{dr} dr ds \end{aligned}$$

This must be counteracted by the gravitational attraction to which the element of mass is subjected. The mass of the element is  $\rho dr ds$ . The attractive force is given by  $\frac{GM(r)\rho dr ds}{r^2}$

For equilibrium

$$-\frac{dP}{dr} dr ds = \frac{GM(r)\rho dr ds}{r^2}$$

Therefore

$$\frac{dP}{dr} = - \frac{GM(r)\rho}{r^2}$$

This is the equation of hydrostatic equilibrium. Integrating from the centre to the surface

$$P(R) - P(0) = - \int_0^R \frac{GM(r)\rho(r)}{r^2} dr$$

But since at the surface  $r = R$ ,  $P(R) = 0$ , the pressure at the centre is

$$P(0) = \int_0^R \frac{GM(r)\rho(r)}{r^2} dr.$$

For uniform density ( $\rho = \text{constant}$ ),

$$M(r) = \frac{4\pi}{3} r^3 \rho$$

$$P(0) = \frac{4\pi}{3} G \rho^2 \int_0^R r dr = \frac{4\pi}{3} G \rho^2 \left[ \frac{r^2}{2} \right]_0^R$$
  
$$= \frac{2\pi}{3} G \rho^2 R^2 = \frac{3}{8\pi} \frac{GM^2}{R^4},$$

where  $M = \frac{4\pi}{3} R^3 \rho$ .

For the Earth  $P(0) \approx 1.7 \times 10^{11} \text{ Pa}$  ( $1.7 \times 10^{12} \text{ dynes/cm}^2$ )  
When a dense core is present as in all planets the estimated value is  $P_{\oplus}(0) \approx 3.6 \times 10^{12} \text{ Pa}$

$$\left\{ \begin{array}{l} (3.6 \times 10^{13} \text{ dynes/cm}^2) \\ (3.6 \text{ MBar}) \end{array} \right.$$